

Instantaneous Spectrum Estimation of Event-Based Densities

Lorenzo Galleani

Dipartimento di Elettronica, Politecnico di Torino, C.so Duca degli Abruzzi 24, 10129 Torino, Italy
Email: galleani@polito.it

Leon Cohen

City University of New York, 695 Park Avenue, New York, NY 10021, USA
Email: leon.cohen@hunter.cuny.edu

Douglas Nelson

US Department of Defense, Fort Meade, MD 20755, USA
Email: waveland@erols.com

Jeffrey D. Scargle

Space Science Division, NASA Ames Research Center, Moffett Field, CA 94035, USA
Email: jeffrey@cosmic.arc.nasa.gov

Received 31 July 2001 and in revised form 30 September 2001

We present a method for obtaining a time-varying spectrum that is particularly suited when the data are in event-based form. This form arises in many areas of science and engineering, and especially in astronomy, where one has photon counting detectors. The method presented consists of three procedures. First, estimating the density using the kernel method; second, highpass filtering the manifestly positive density; finally, obtaining the time-frequency distribution with a modified Welch's method. For the sake of validation, event-based data are generated from a given distribution and the proposed method is used to construct the time-frequency spectrum and is compared to the original density. The results demonstrate the effectiveness of the method.

Keywords and phrases: time-frequency analysis, density estimation, kernel method.

1. INTRODUCTION

An important problem in many fields is the study of densities and spectra where the data are event-based. We illustrate with a specific example of the study of a signal coming from the X-ray binary system XTE-J1550. The signal is made by a sequence t_i of the photon arrival times, measured by an X-ray detector orbiting the Earth [1, 2]. The more concentrated are the photons in a given time interval, the higher will be the intensity of the radiation (in the X-ray region) for that particular time interval. Hence the density $f(t)$ of the sequence t_i is here intended as the normalized intensity of the radiation underlying the arriving photons.

Since a fundamental property of such a density is its instantaneous spectrum [3, 4], because it is believed to carry important physical information on the system that generated the events, we present here a method that enables us to perform this estimation in an effective way. We want to build an

estimation algorithm that performs the following operation

$$t_i \longrightarrow P(t, \omega), \quad (1)$$

where $P(t, \omega)$ is the estimated instantaneous spectrum of the (normalized) density $f(t)$ underlying the sequence of events t_i . The astronomy case mentioned earlier and described in detail in [1, 2] can be considered as an excellent paradigm for instantaneous spectrum estimation of event-based densities. Here the main physical quantity is time, but the identical method can be applied to any quantity x (with associated density $f(x)$).

The method consists of three steps. First, a density estimation is accomplished by using the kernel method, namely sliding a window that makes a weighted count of the events in the window support. Then highpass filtering allows one to eliminate the low frequency components that are always present in a density due to its positivity, but are often of no

interest. Finally, the estimation of the time-frequency spectrum is performed by using a sliding Welch's estimator to recover instantaneous spectral components even in very noisy environments. The method is tested on simulated event-based data, and the results show its robustness and effectiveness in estimating the time-varying spectral features of the initial density $f(t)$.

2. METHOD

We now describe the basic steps of the method for the instantaneous spectrum estimation of event-based densities.

2.1. Density estimation

The original data represent a sequence of events t_i that is typically recorded as a sequence of the time instants when the events occurred. This sequence of events has a density $f(t)$. In the physical example that we mentioned in the introduction, the events are the photon arrival times in the X-ray detector mounted on the satellite, and the density is proportional to the strength of the incident radiation field. If no prior information is given on the stochastic process that underlies the events, the first step is to estimate the density $f(t)$ underlying the data.

To accomplish such estimation we use the kernel method [5]. This method consists in sliding a function $h(t)$ called kernel on the event data t_i , and then operating a summation of the events in the kernel support using the kernel itself as weighting function. The kernel must be a positive function with integral equal to one. The estimated density $\hat{f}(t)$ at a certain value t is the weighted summation obtained when the kernel is centered at that analysis value t . The kernel method generates a continuous estimated density. The shape and analytic properties of the kernel can control the properties of the estimated density. For example, differentiability of $\hat{f}(t)$ holds if the kernel is differentiable.

In [1, 2] we used a Hanning kernel, whereas in this paper results are obtained using a rectangular kernel. In general, the kernel must be chosen according to the nature of the data.¹ Also, in [7] a filter design approach has been discussed for kernel choice. Some general rules that we can follow in choosing an appropriate kernel are:

- If we want the estimated density $\hat{f}(t)$ to have certain analytical properties, these properties must be imposed on the kernel $h(t)$.
- As is the case in selecting a window for generating a spectrogram, tracking of fast variations in $f(t)$ is better accomplished by more concentrated kernels, such as rectangular kernels.
- The kernel method can be written as a convolution where the kernel $h(t)$ plays the role of a filter [7]. In this sense one can choose the kernel by noting the effect on

the spectrum of the estimated density (noise reduction, resolution, etc.).

2.2. Highpass filtering

The estimated density $\hat{f}(t)$ is a positive function, and hence its spectrum always shows strong low frequency components and so will the time-frequency spectrum. Since the spectral components near zero are often not of interest, we highpass filter $\hat{f}(t)$ to remove them. The digital filter used in the operation can be a standard one, but care must be taken to not introduce nonlinear phase distortion in the filter process. This is particularly important in the time-frequency plane, where instantaneous spectral components are connected to the derivative of the instantaneous phase, and hence phase distortion directly changes the instantaneous spectrum. A forward-backward filtering is chosen, that consists in filtering, reversing, and filtering again the signal. The output density $\hat{f}_h(t)$ can be proven to have zero phase distortion [8].

2.3. Time-frequency estimation

The last step is the estimation of the instantaneous spectrum $P(t, \omega)$ from the estimated density $\hat{f}_h(t)$ obtained from the previous step. Since the density is always corrupted with noise, we have to adopt noise reduction techniques. This is accomplished by using the sliding estimator, a time-frequency estimation technique [9, 10, 11, 12] (see also [13]). The sliding estimator is simply obtained by sliding a window on the signal centered at every analysis time, and then taking the Welch's periodogram of the windowed signal at that time. The result is the extension of the well-known periodogram to the non-deterministic case of the time-frequency spectrum estimation of a signal embedded in noise.

3. VALIDATION OF THE METHOD

To prove the validity of the method, we test it on event-based densities generated by numerical algorithms. Since the generation of events with a fixed density is not a common problem in signal analysis, we provide a description of the method adopted for such purpose. After that we present and discuss the results of the time-frequency estimation method applied to the generated densities. We point out that our method has been motivated primarily by timing problems arising in astronomy as discussed in the introduction [1, 2, 14]. The only current method (to the best of our knowledge) that allows time-frequency spectrum estimation of event-based densities is the so-called "Dynamic Power Spectrum" [14, 15]. We have already discussed this method and shown comparisons with our algorithm in [1, 2] and hence we will not repeat them here. Our method shows an improvement in the details of the time-varying spectrum.

Before describing how to generate events, we first choose a continuous density $f(t)$ by fixing a signal with a given time-frequency law. In particular, we generate a signal that has a time-frequency behavior similar to the QPO (Quasi-Periodic Oscillation, see [16]) that we investigated in [1, 2]. In Figure 1, we show the time-frequency spectrum of the chosen density

¹The issue of what is choosing a kernel has a long history. See [6] for a discussion of this issue.

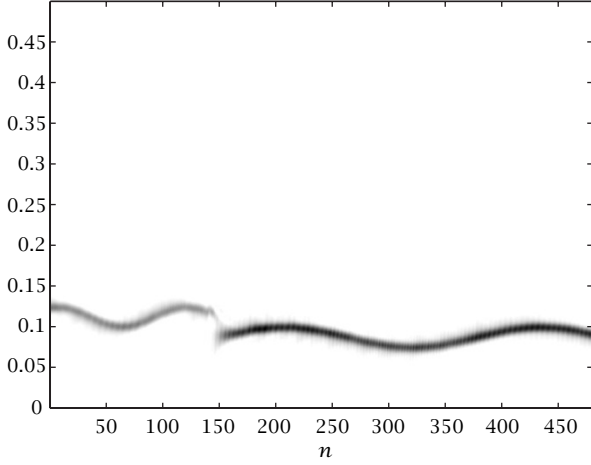


FIGURE 1: Time-frequency spectrum of a regularly sampled version of $f(t)$. The signal is similar to the QPO (Quasi-Periodic Oscillation) presented in [1, 2].

$f(t)$ obtained with the sliding estimator technique described in Section 2.3. The density in this example is still regularly sampled as a standard digital signal. A better time-frequency distribution could be used to represent the signal, for example a Choi-Williams distribution [17] that exhibits greater resolution, but here we use the sliding estimator because we will compare against Figure 1 all the obtained results, and so we need to compute them in a similar way.

3.1. Density generation

We now describe a method to generate a series of events, t_i , that fit the chosen density $f(t)$. These, for example, may be time of arrival events. Among the several methods available, we choose the mixing method [18], for its simplicity and reliability. The mixing method is based on the possibility of writing the desired distribution $f(t)$ as a linear combination of distributions, that is,

$$f(t) = p_1 f_1(t) + p_2 f_2(t) + \dots + p_n f_n(t), \quad p_k > 0, \quad (2)$$

where the constants p_1, p_2, \dots, p_n play the role of weights and, due to the density property of all the functions f_k involved, also probabilities. We generate the sequence t_i by mixing m sequences t_i^k that are selected in the following way:

$$\text{set } t_i = t_i^k \quad \text{if } p_1 + \dots + p_{k-1} \leq u_i < p_1 + \dots + p_k, \quad (3)$$

where u_i is a sequence with uniform distribution between 0 and 1. This decomposition is useful if the basic densities $f_k(t)$ are easier to generate by numerical algorithms. The easiest case is, of course, when we choose the f_k to be uniform distributions. By applying the mixing method one can approximate any given distribution to any desired degree of accuracy.

In particular, we choose the f_k densities to be linear densities. In this case we can approximate $f(t)$ with greater convergence speed than by using uniform distributions. Hence

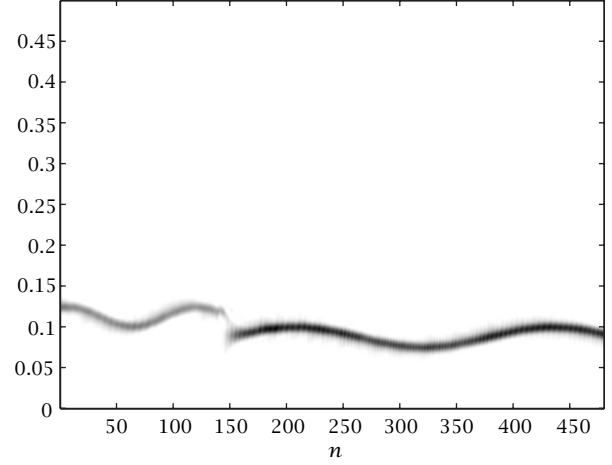


FIGURE 2: Estimated instantaneous spectrum of the density $f(t)$ when a sequence t_i of events is given. The number of events is twice the number of samples of the numerical version of $f(t)$.

we approximate $f(t)$ using $N - 1$ linear densities having supports in the time intervals $[kTs, (k+1)Ts), k = 0, \dots, N - 2$, where N is the number of samples of $f(t)$ (we consider its sampled version). The weights p_1, p_2, \dots, p_n are computed as the $N - 1$ areas of $f(t)$ in the $N - 1$ time supports (their sum equals one because $f(t)$ is a proper density).

3.2. Estimation results

We now present the time-frequency spectrum obtained applying the proposed method to the chosen density $f(t)$. We consider the noise free case and then we add to the original density a white Gaussian noise with decreasing SNR, to simulate measurement noise and other stochastic processes embedded in the data. We point out that often in astronomy the only source of noise are the statistical fluctuations associated with counting a finite number of discrete events. Nevertheless, in the general framework of signal analysis, it is a basic and important approach to consider the case where additive white noise is present.

Noise-free case

In Figure 2 we represent the estimated instantaneous spectrum of $f(t)$ obtained from a sequence t_i of events generated by the mixing method. The number of events used is $2N$, N being the length of the original sampled density $f(t)$. The quality of the result is surprising, if one considers that only two events are generated for each of the N samples of the noise free density. To give an idea on how rough is the approximation of $f(t)$ obtained using $2N$, we plot in Figure 3 a comparison between the original density $f(t)$ and a reconstructed version $\hat{f}(t)$ obtained from the kernel method. The quality of the reconstruction increases if one generates a sequence t_i made of $4N$ events, as Figure 4 shows. Here a direct comparison with Figure 1 shows the quality of the estimation.

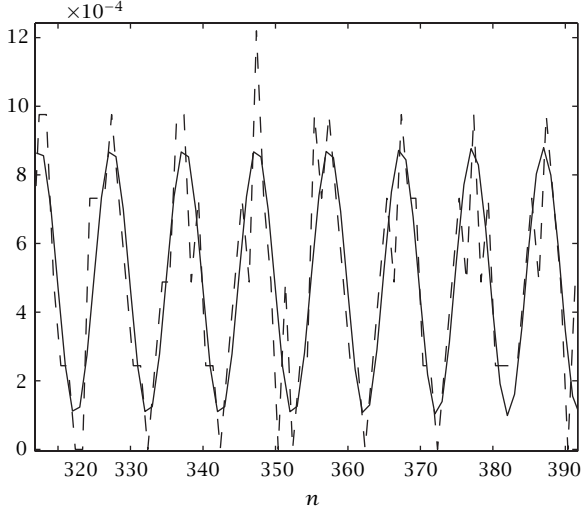


FIGURE 3: Comparison between the initial density $f(t)$ (continuous line) and the estimated version $\hat{f}(t)$ (dashed line) obtained with the kernel method. Notice how poor seems the quality of the estimated density $\hat{f}(t)$. Despite this apparently bad estimation, the estimated instantaneous power spectrum will show clearly the presence of a time-varying frequency in the data.

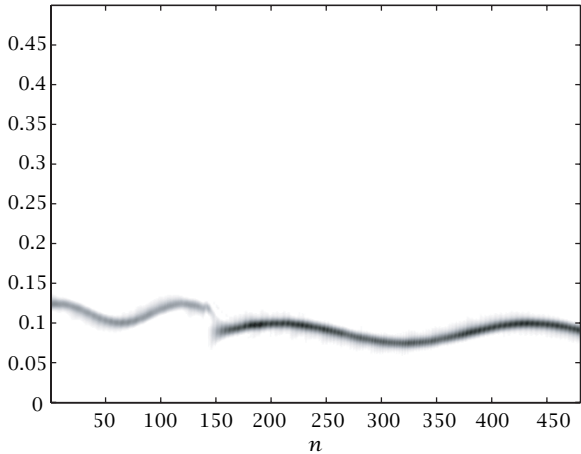


FIGURE 4: Estimated instantaneous spectrum of the density $f(t)$ when a sequence t_i of events is given. The number of events is four times the number of samples of the numerical version of $f(t)$.

SNR = 2.5 dB

In Figure 5 we represent the estimation done when SNR = 2.5 dB. The reconstruction is very good, despite only $4N$ events have been used to generate the sequence t_i .

SNR = -7 dB

In Figure 6 we plot the estimation obtained when SNR = -7 dB and $40N$ events are used. The noise has now increased consistently, and $4N$ points are not enough to retrieve the time-frequency signature of the original signal. The increased number of events allows us to clearly understand the structure

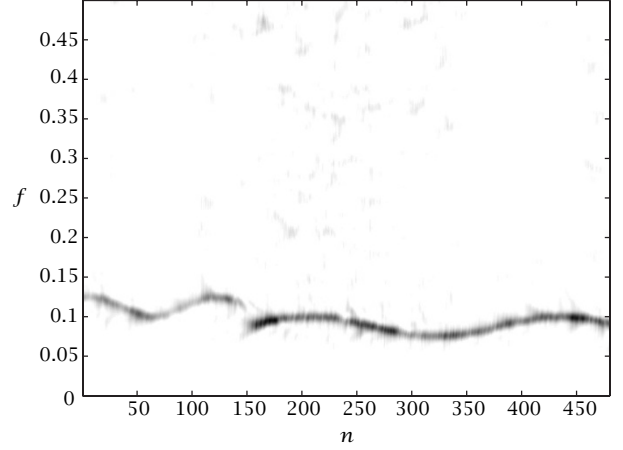


FIGURE 5: Time-frequency spectrum estimated when SNR = 2.5 dB. Notice the quality of the approximation, even if only $4N$ events have been used (N being the length of the numerical version of $f(t)$).

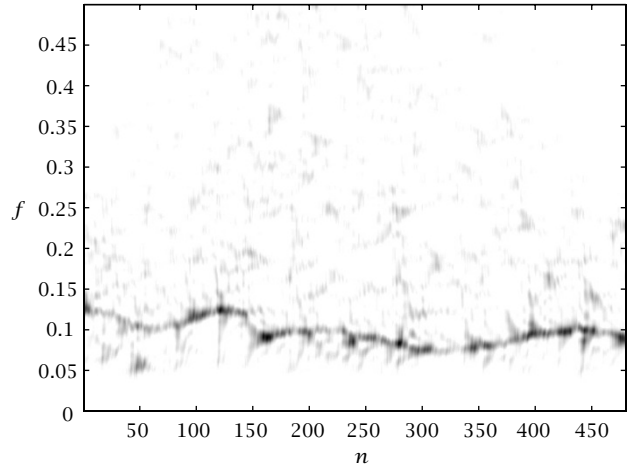


FIGURE 6: Time-frequency spectrum estimated with the proposed method when SNR = -7 dB. The number of events is $40N$, to guarantee a sufficient estimation of the original density.

of the time-varying frequency component. Also a “bumpy” behavior is noticeable in the signal component. This is due mainly to the interaction between the high noise level and the signal itself, more than to the event-based nature of the data. The latter can be in fact excluded considering the results obtained in the noise free case, where the event-based time-frequency spectrum represents an excellent approximation of the original density, with no visible “bump.”

4. CONCLUSION

We have presented a new method that allows the estimation of the instantaneous spectrum of a density $f(t)$ when only a sequence of events t_i (with density $f(t)$) is given. We considered the case of a “noisy density,” that means we developed the estimation algorithm taking care of noise reduction by

using a Welch's algorithm. The results presented show that the estimation algorithm is robust and effective, and that has excellent performance even when the number of events t_i available is small and a high noise level is present.

REFERENCES

- [1] L. Galleani, L. Cohen, D. Nelson, and J. Scargle, "Time-evolution of the power spectrum of the black hole X-ray nova XTE J1550-564," in *Nonlinear Signal and Image Processing (NSIP-01)*, Baltimore, USA, June 2001.
- [2] L. Galleani, L. Cohen, D. Nelson, and J. Scargle, "Estimating the instantaneous power spectrum of an X-ray binary system," in *SPIE 2001*, San Diego, USA, 30 July-3 August 2001.
- [3] L. Cohen, "Time-frequency distributions—a review," *Proceedings of the IEEE*, vol. 77, no. 7, pp. 941–981, 1989.
- [4] L. Cohen, *Time-Frequency Analysis*, Prentice-Hall, New Jersey, 1995.
- [5] B. W. Silverman, *Density Estimation for Statistics and Data Analysis*, Chapman & Hall/CRC Press, New York, January 1986.
- [6] A. Papoulis, *Signal Analysis*, McGraw-Hill, New York, 1984.
- [7] D. W. Scott, *Multivariate Density Estimation*, John Wiley & Sons, New York, 1992.
- [8] A. Antoniou, *Digital Filters: Analysis and Design*, McGraw-Hill, New York, 1979.
- [9] G. Frazer and B. Boashash, "Multiple window spectrogram and time frequency distributions," in *Proc. of the IEEE Int. Conference on Acoustics, Speech and Signal Processing (ICASSP-94)*, vol. 6, pp. 293–296, Adelaide, Australia, 1994.
- [10] M. Bayram and R. G. Baraniuk, "Multiple window time-frequency analysis," in *Proc. of the IEEE-SP Int. Sym. on Time-Frequency and Time-Scale Analysis*, pp. 173–176, 1996.
- [11] J. Pitton, "Time-frequency spectrum estimation: an adaptive multitaper method," in *IEEE Int. Sym. Time-Frequency and Time-Scale Analysis*, pp. 665–668, 1998.
- [12] F. Cakrak and P. Loughlin, "Multiple window non-linear time-varying spectral analysis," *IEEE Trans. Signal Processing*, vol. 49, no. 2, pp. 448–453, 2001.
- [13] D. Nelson, "Special purpose correlation functions for improved signal detection and parameter estimation," in *Proc. of IEEE Conference on Acoustics, Speech and Signal Processing*, pp. 73–76, April 1993.
- [14] J. Homan, R. Wijnands, M. van der Klis, et al., "Correlated X-ray spectral and timing behavior of the black hole candidate XTE J1550-564: a new interpretation of black hole states," *Astrophysical Journal Supplement Series*, vol. 132, pp. 377–402, February 2001.
- [15] M. van der Klis, "Fourier techniques in X-ray timing," in *Proc. NATO Advanced Study Institute, Timing Neutron Stars*, pp. 27–69, 1989.
- [16] J. Scargle, D. L. Donoho, J. P. Crutchfield, T. Steiman-Cameron, J. Imamura, and K. Young, "The quasi-periodic oscillations and very low frequency noise of scorpius X-1 as transient chaos: a dripping handrail?," *Astrophys. J. Lett.*, vol. 411, pp. L91–L94, 1993.
- [17] W. Williams, "Reduced interference distributions: biological applications and interpretations," *Proceedings of the IEEE*, vol. 84, pp. 1264–1280, 1996.
- [18] A. Papoulis, *Probability, Random Variables, and Stochastic Processes*, McGraw-Hill, 3rd edition, 1991.

Lorenzo Galleani was born in 1970 in Torino, Italy. He received the B.S. and Ph.D. degrees in Electrical Engineering from Politecnico di Torino, in 1997 and 2001, respectively. He is a postdoctoral researcher at Hunter College, City University of New York, and at Politecnico di Torino. His main research interests are on modern spectral analysis and dynamical systems.



Leon Cohen received the B.S. degree from City College in 1962 and the Ph.D. degree from Yale University in 1966. He is currently Professor of Physics at Hunter College and Graduate Center of the City University of New York. He has done research in mathematical physics, astronomy, quantum mechanics, chemical physics, and signal analysis. He has held research visiting positions at Harvard-Smithsonian Center for Astrophysics, University of North Carolina, Center for Advanced Studies at the University of New Mexico, IBM T. J. Watson Research Center and Naval Underwater Systems Center. He is the author of the recently published book "Time-Frequency Analysis" (Prentice Hall).



Douglas Nelson was born in Minneapolis, Minnesota in 1945. He received a BA in Mathematics from the University of Minnesota in 1967 and his Ph.D. in Mathematics from Stanford University in 1972. He was an assistant professor of Mathematics at Carnegie-Mellon University until 1975, when he started working at the National Security Agency, where he has been from 1975 to the present. At the National Security Agency, Dr Nelson has worked primarily on the development and implementation of signal processing algorithms for a variety of applications, including radar analysis, communications and speech.



Jeff Scargle received his B.A., Summa Cum Laude from Pomona College in 1963 and Ph.D. from California Institute of Technology in 1968. He has been with Space Science Division of NASA Ames Research Center since 1975. His research interests are high energy and planetary astrophysics, with specialization on gamma-ray and X-ray astronomy, time series analysis, and nonlinear dynamical systems and chaos.

